

# Northwest Africa 4898

Unbrecciated basalt

137 g



Figure 1: Slice through NWA 4898, with close up of matrix (upper right) and 1 cm cube for scale.

## **Introduction**

Northwest Africa 4898 (Fig. 1) was found in northwest Africa in 2007, and consists of one fragment almost completely covered with fusion crust and weighing 137 g (Connolly et al., 2008). Inspection of the interior reveals its basaltic texture (Fig. 1).

## **Petrography and Mineralogy**

The texture of this sample is spherulitic with lath-shaped plagioclase, pyroxene, and skeletal ilmenite. Olivine is present as larger crystals ( $\text{Fa}_{26.3-27.2}$ ;  $\text{FeO/MnO} = 73-92$ ), and often contains chromite inclusions. The calcic plagioclase ( $\text{An}_{92.6-96.5}$ ) has been completely transformed into maskelynite during shock metamorphism. And the pyroxenes ( $\text{Fs}_{25.1-58.7}\text{Wo}_{13.2-34}$ ;  $\text{FeO/MnO} = 42-76$ ) are compositionally zoned Ti-rich pigeonite and augite. FeNi-metal and troilite are present as minor phases (Connolly et al., 2008).

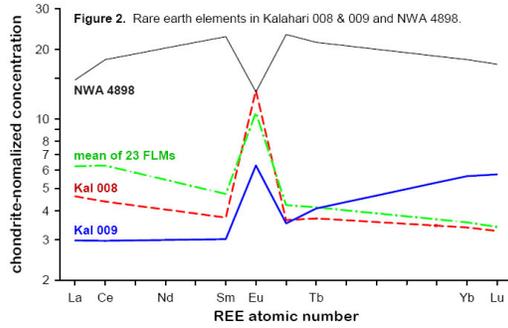


Figure 2: Rare earth element diagram for NWA 4898 illustrating a Eu anomaly (from Korotev et al., 2008).

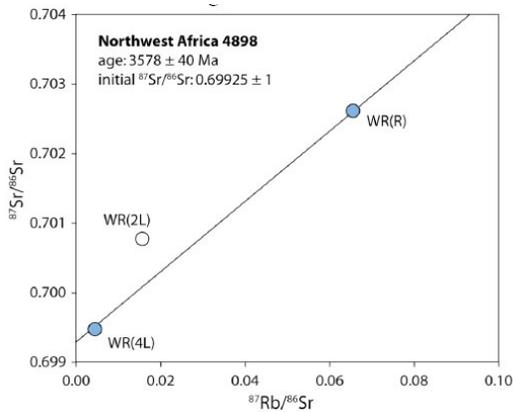


Figure 3: Whole rock isochron for NWA4898 indicating an age of 3.58 Ga (from Gaffney et al., 2008).

## Chemistry and Radiogenic age dating

INAA analyses of 6 small (133 mg total) chips of NWA 4898 reveal its low FeO nature (Table 1) and Eu anomaly (Fig. 2). A whole rock Rb-Sr isochron based on three measurements from NWA 4898 yields an age of 3.58 Ga (Fig. 3; Gaffney et al., 2008). Additional Nd isotopic analyses show that NWA 4898 may have been derived from one of the most LREE depleted lunar mantle sources known (Fig. 4; Gaffney et al., 2008).

## Cosmogenic isotopes and exposure ages

None yet reported.

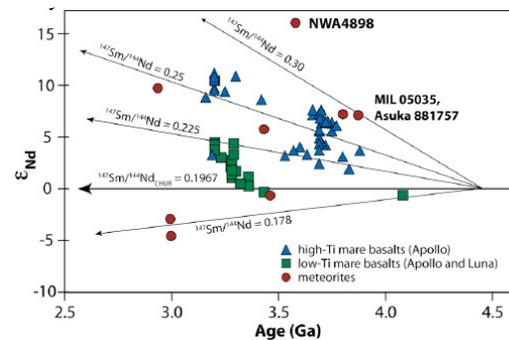


Figure 4: Results of Nd isotopic analyses of Gaffney et al (2008) showing the depleted nature of this sample relative to other Apollo, Luna and meteorite samples.

**Table 1. Chemical composition of NWA 4898**

<i>reference</i>	1	Rh	
<i>weight</i>	133	Pd ppb	
<i>technique</i>	INAA	Ag ppb	
		Cd ppb	
		In ppb	
SiO <sub>2</sub> %	-	Sn ppb	
TiO <sub>2</sub>	-	Sb ppb	
Al <sub>2</sub> O <sub>3</sub>	-	Te ppb	
Cr <sub>2</sub> O <sub>3</sub>	-	Cs ppm	
FeO	17.2	Ba	
MnO	-	La	
MgO	-	Ce	
CaO	-	Pr	
Na <sub>2</sub> O	0.296	Nd	
K <sub>2</sub> O	-	Sm	4.55
P <sub>2</sub> O <sub>5</sub>	-	Eu	0.997
S %	-	Gd	
<i>sum</i>	-	Tb	
		Dy	
Sc ppm	65.4	Ho	
V		Er	
Cr	3020	Tm	
Co		Yb	
Ni	<180	Lu	
Cu		Hf	
Zn		Ta	
Ga		W ppb	
Ge		Re ppb	
As		Os ppb	
Se		Ir ppb	
Rb		Pt ppb	
Sr		Au ppb	
Y		Th ppm	0.44
Zr		U ppm	
Nb			
Mo			
Ru			

References: 1) Korotev et al. (2008).